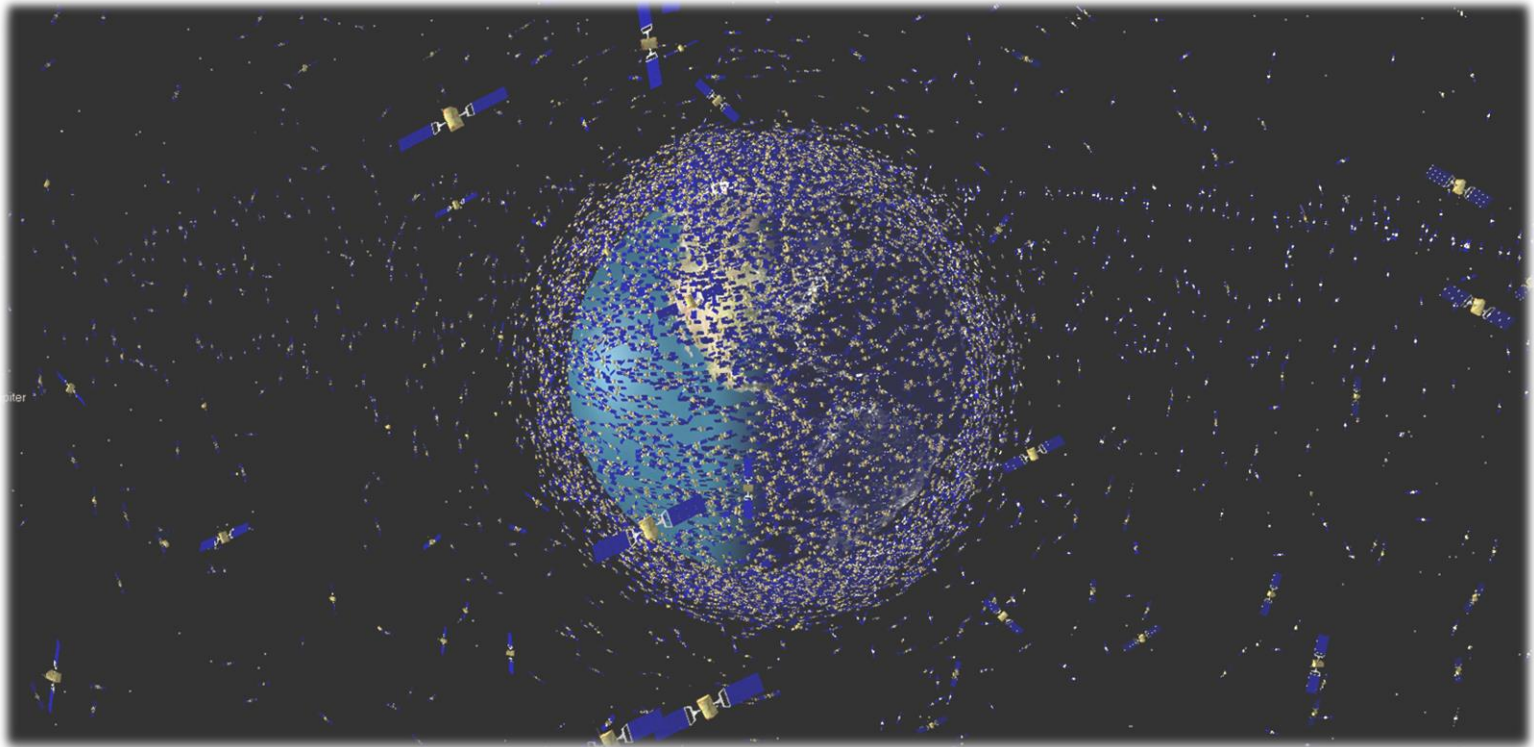
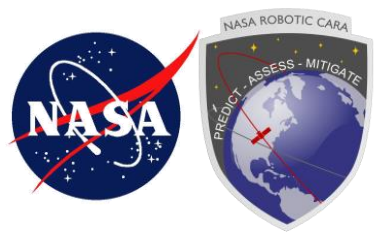




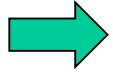
# NASA Robotic CARA Satellite State Estimate Covariance



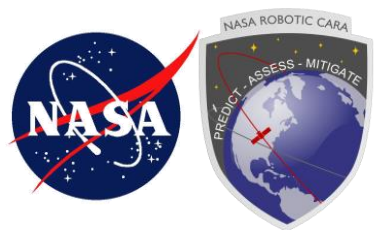
**CARA Operations Team**  
**[Cara-ops@lists.nasa.gov](mailto:Cara-ops@lists.nasa.gov)**  
**301-789-4306**



# Agenda



- **Covariance basics**
- **Use of covariance in probability of collision ( $P_c$ ) calculation**
- **Covariance generation and propagation methods**
- **Covariance tuning**
- **Covariance realism improvement methods**
- **Covariance theory compatibility**

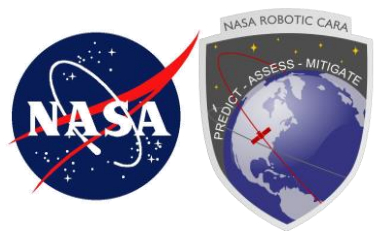


- **Purpose of OD**

- Generate estimate of the object's state at a given time (called the *epoch time*)
- Generate additional parameters and constructs to allow object's future states to be predicted (accomplished through orbit *propagation*)
- Generate a statement of the estimation error, both at epoch and for any predicted state (usually accomplished by means of a *covariance matrix*)

- **Error types**

- OD approaches (either batch or sequential estimation) presume that they solve for all significant systematic errors
- Remaining solution error is thus presumed to be random (Gaussian) error
- Sometimes estimates of this error can be intentionally inflated to try to improve the fidelity of the error modeling
- Nonetheless, presumed to be Gaussian in form and unbiased
  - Issue of non-Gaussian error volumes addressed later in package



# OD Parameters Generated by ASW Solutions

- **Solved for: State parameters**

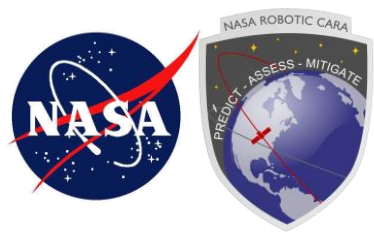
- Six parameters needed to determine 3-d state fully
- Cartesian: three position and three velocity parameters in orthogonal system
- Element: six orbital elements that describe the geometry of, and object location within, the orbit

- **Solved for: Non-conservative force parameters**

- Ballistic coefficient ( $C_D A/m$ ); describes susceptibility of spacecraft state to atmospheric drag
- Solar radiation pressure (SRP) coefficient ( $C_R A/m$ ); describes susceptibility of spacecraft state to radiation momentum from sun

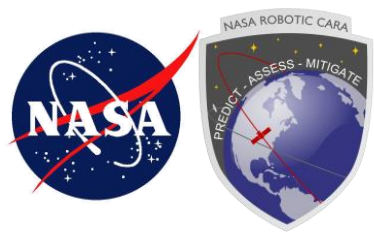
- **Considered: ballistic coefficient and SRP consider parameter**

- Not solved for but “considered” as part of the solution
- Derived from information outside of the OD itself
- Discussed later



# OD Uncertainty Modeling

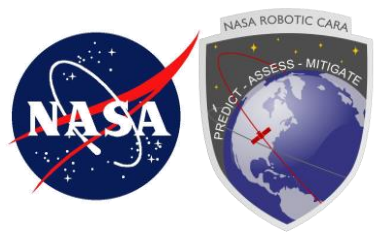
- **Characterizes the overall uncertainty of the OD epoch and/or propagated state**
  - Uncertainty of each estimated parameter and their interactions
- **This is a characterization of a multivariate statistical distribution**
- **In general, need more than two moments to describe a distribution**
  - Mean, variance, skewness, and kurtosis often used
  - Requires higher-order tensors to do this for a multivariate distribution
- **Assumptions about error distribution can simplify situation substantially**
  - Presuming the solution is unbiased places the mean error values at zero
  - Presuming the error distribution is Gaussian eliminates the need for the third and fourth moments (skewness and kurtosis)
  - Error distribution can thus be expressed by means of variances of each solved-for component and their cross-correlations
  - Thus, error can be fully represented by means of a covariance matrix



# Covariance Matrix Construction: Symbolic Example

- Three estimated parameters (a, b, and c)
- Variances of each along diagonal
- Off-diagonal terms the product of two standard deviations and the correlation coefficient ( $\rho$ ); matrix is symmetric

	a	b	c	...
a	$\sigma_a^2$	$\rho_{ab}\sigma_a\sigma_b$	$\rho_{ac}\sigma_a\sigma_c$	...
b	$\rho_{ab}\sigma_a\sigma_b$	$\sigma_b^2$	$\rho_{bc}\sigma_b\sigma_c$	...
c	$\rho_{ac}\sigma_a\sigma_c$	$\rho_{bc}\sigma_b\sigma_c$	$\sigma_c^2$	...
...	...	...	...	...



# Example Covariance from CDM

- **8 x 8 matrix typical of most ASW updates**

- Some orbit regimes not suited to solution for both drag and SRP; these covariances 7 x 7

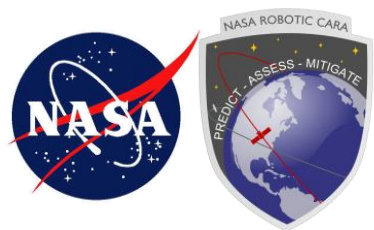
- **Mix of different units often creates poorly conditioned matrices**

- Condition number of matrix at right is  $9.8E+11$ —terrible!

- **Often better numerically (and more intuitive) to separate matrix into sections**

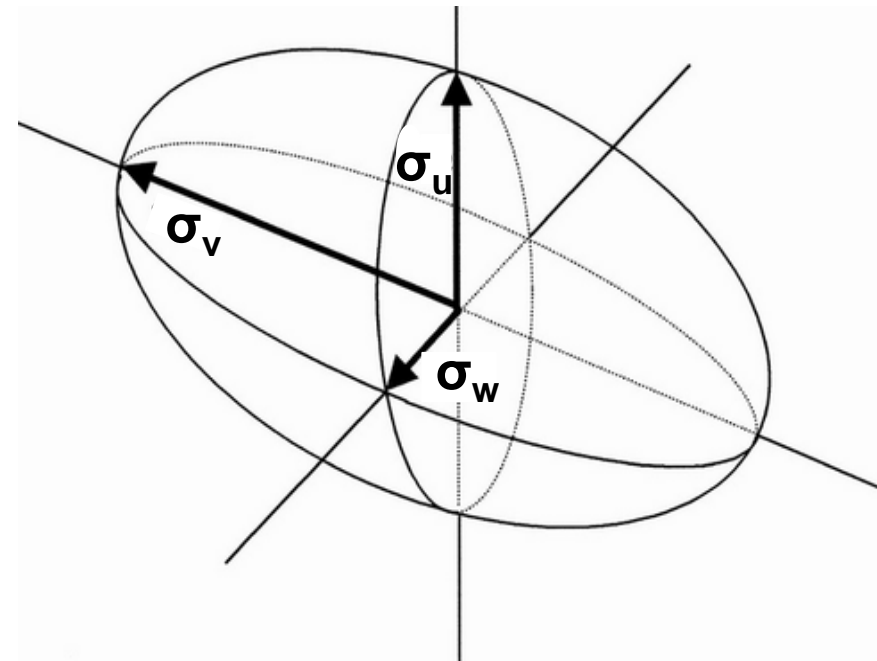
- **First 3 x 3 portion (amber) is *position* covariance—often considered separately**

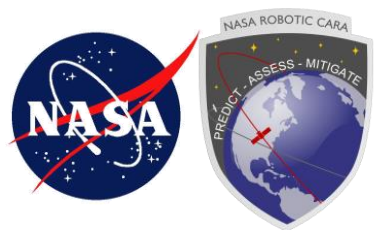
	U (m)	V (m)	W (m)	Udot (m/s)	Vdot (m/s)	Wdot (m/s)	B (m <sup>2</sup> /kg)	AGOM (m <sup>2</sup> /kg)
U	6.84E+01	-2.73E+02	6.38E+00	2.76E-01	-7.14E-02	8.75E-03	-3.83E-02	-3.83E-02
V	-2.73E+02	1.10E+05	3.23E+01	-1.17E+02	-8.99E-02	2.51E-02	-1.28E-01	-1.28E-01
W	6.38E+00	3.23E+01	4.47E+00	-3.26E-02	-6.83E-03	1.81E-03	-3.73E-03	-3.73E-03
Udot	2.76E-01	-1.17E+02	-3.26E-02	1.24E-01	1.10E-04	-2.47E-05	1.46E-04	1.46E-04
Vdot	-7.14E-02	-8.99E-02	-6.83E-03	1.10E-04	7.57E-05	-9.39E-06	4.10E-05	4.10E-05
Wdot	8.75E-03	2.51E-02	1.81E-03	-2.47E-05	-9.39E-06	2.06E-05	-4.39E-06	-4.39E-06
B	-5.07E-03	1.30E+00	4.34E-05	-1.38E-03	7.97E-07	7.26E-07	1.64E-05	-6.28E-07
AGOM	-3.83E-02	-1.28E-01	-3.73E-03	1.46E-04	4.10E-05	-4.39E-06	-6.28E-07	2.31E-05



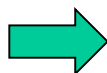
# Position Covariance Ellipse

- **Position covariance defines an “error ellipsoid”**
  - Placed at predicted satellite position
  - Square root of variance in each direction defines each semi-major axis (UVW system used here)
  - Off-diagonal terms rotate the ellipse from the nominal position shown
- **Ellipse of a certain “sigma” value contains a given percentage of the expected data points**
  - 1- $\sigma$ : 19.9%
  - 2- $\sigma$ : 73.9%
  - 3- $\sigma$ : 97.1%
  - Note how much lower these are than the univariate normal percentage points

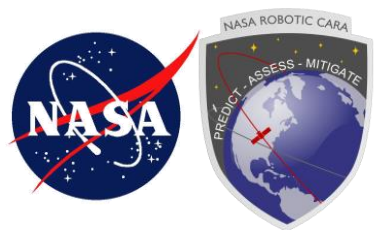




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- Covariance basics
- Use of covariance in probability of collision ( $P_c$ ) calculation
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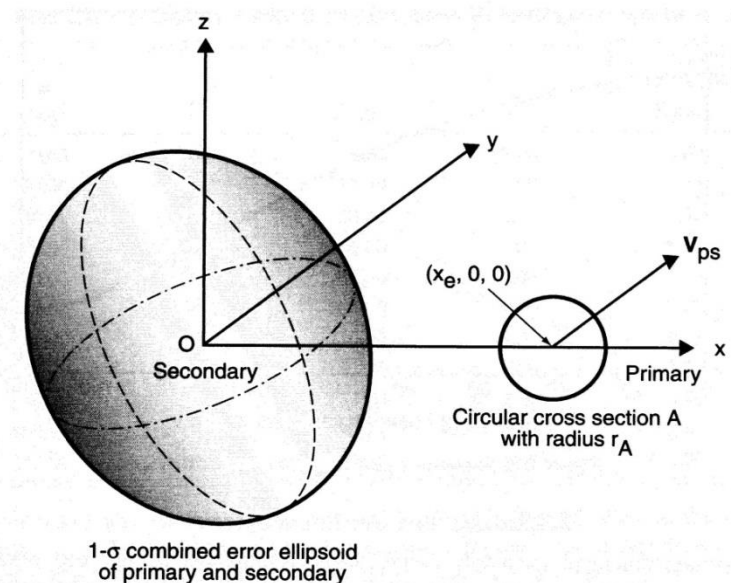


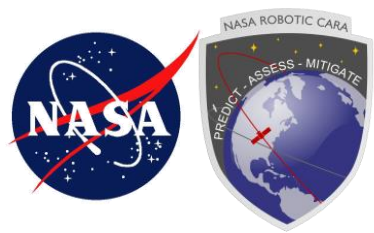
# Covariance in Calculation of 2D Probability of Collision ( $P_c$ )

- Primary and secondary covariances combined and projected into 2D *conjunction plane* (plane perpendicular to velocity vector)
- Primary placed on x-axis at (miss distance, 0) and represented by circle of radius equal to sum of both spacecraft circumscribing radii
- Z-axis perpendicular to x-axis in conjunction plane
- $P_c$  is portion of combined error ellipsoid that falls within the hard-body radius circle

$$P_c = \frac{1}{\sqrt{(2\pi)^2 |C^*|}} \iint_A \exp\left(-\frac{1}{2} \mathbf{r}^T C^{*-1} \mathbf{r}\right) dX dZ$$

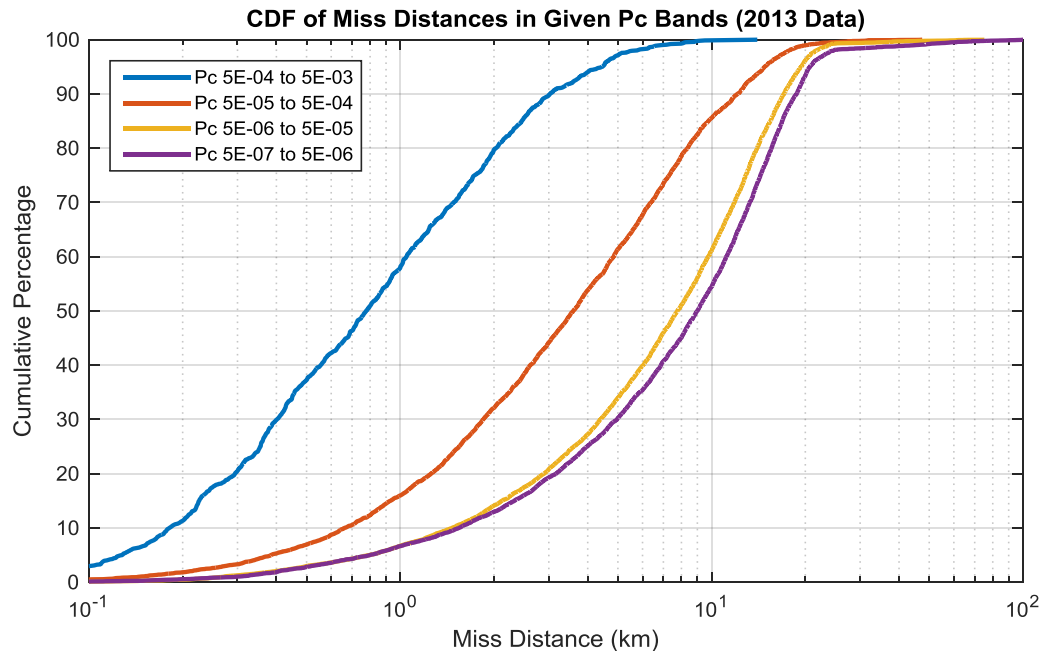
**Covariance essential to  $P_c$  calculation, which is the most important factor in collision risk assessment**

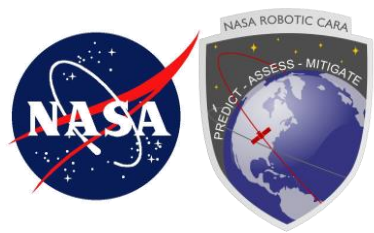




# Pc vs Miss Distance Calculations

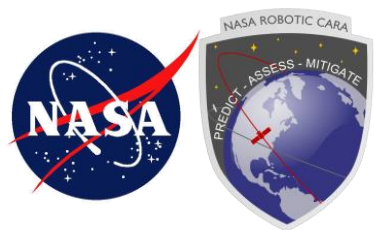
- **Pc is the best single-parameter encapsulation of the risk**
- **Without Pc, have only the miss distance**
- **Correlation between miss distance and Pc very poor**
  - Four Pc bands shown below; correlation with miss distance poor in all cases
- **Important to have Pc, and covariance necessary for its calculation**





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# Batch Epoch Covariance Generation

- **Batch least-squares update (ASW method) uses the following minimization equation**

- $dx = (A^TWA)^{-1}A^TWb$

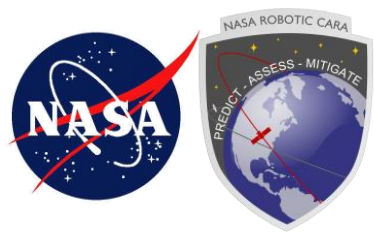
- $dx$  is the vector of corrections to the state estimate
    - $A$  is the time-enabled partial derivative matrix, used to map the residuals into state-space
    - $W$  is the “weighting” matrix that provides relative weights of observation quality (usually  $1/\sigma$ , where  $\sigma$  is the standard deviation generated by the sensor calibration process)
    - $b$  is the vector of residuals (observations – predictions from existing state estimate)

- **Covariance is the collected term  $(A^TWA)^{-1}$**

- **$A$**  the product of two partial derivative matrices:

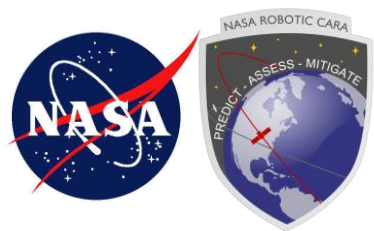
- $A = \frac{\partial(obs)}{\partial X_0} = \frac{\partial(obs)}{\partial X} \frac{\partial X}{\partial X_0}$

- First term: partial derivatives of observations with respect to state at obs time
      - Second term: partial derivatives of state at obs time with respect to epoch state



# Batch Epoch Covariance Generation

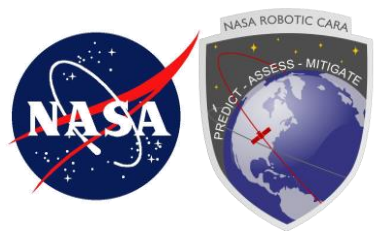
- **Formulated this way, this covariance matrix is called an *a priori* covariance**
  - **A** does not contain actual residuals, only transformational partial derivatives
  - So  $(A^TWA)^{-1}$  is a function only of the amount of tracking, times of tracks, and sensor calibration relative weights among those tracks
    - Not a function of the actual residuals from the correction
    - Rather, an *a priori* estimate of the errors that should be experienced given the tracking density, times, and relative accuracies
- **Limitations of *a priori* covariance**
  - Does not account well for unmodeled errors, such as transient atmospheric density prediction errors
    - Because not examining actual fit residuals
  - W-matrix only as good as sensor calibration process
    - Principal weakness of present process



# Covariance Propagation

## Method 1: Full Monte Carlo

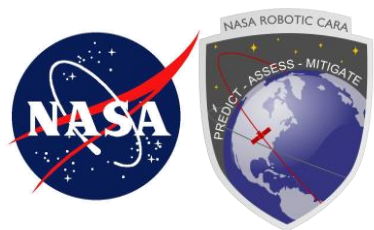
- **Creates  $n$  state (position and velocity) perturbations at epoch**
  - Covariance at epoch describes uncertainty of state at epoch
  - Can use this to create set of  $n$  possible realizations of the epoch state, conforming to the distribution parameters specified by the covariance
- **Propagates each of these forward to the time of interest**
  - Use the full non-linear dynamics of the propagator
  - Thus produce  $n$  states at TCA (for CA application)
- **Summarizes set of  $n$  states statistically**
  - Usually empirically, through non-parametric techniques (e.g., percentiles, empirical distribution functions)



# Covariance Propagation

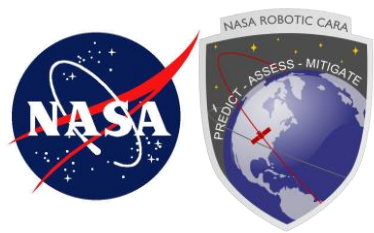
## Method 2: Linear Mapping

- **Non-linear dynamics of orbit propagation can be linearized**
  - These linear approximations valid for “short” periods about epoch state
- **State transition matrix ( $\Phi$ ) the encapsulation of this linearization**
  - Can be used for state propagations (but often is not)
  - Can also be used for propagation of covariance [ $\Phi(t, t_0) * C(t_0) * \Phi^T(t, t_0)$ ]
- **Covariance propagation can also be augmented via the addition of “process noise”**
  - Process noise matrix ( $Q$ ) formulated, which specifies acceleration uncertainty in each coordinate principal direction
    - Intent is to compensate for unmodeled and inadequately-modeled perturbations
    - Can potentially remediate some of the limitations introduced by the linearization
  - Process noise matrix propagated through use of a process noise transition matrix, in a manner similar to state transition: [ $\Gamma(t, t_0) * Q(t_0) * \Gamma^T(t, t_0)$ ]



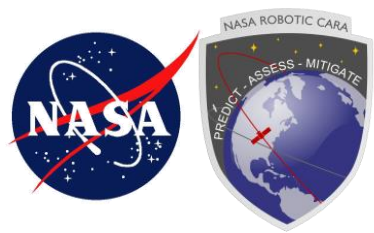
# Covariance Propagation Method Comparison

Method	Advantages	Disadvantages
<b>Monte Carlo</b>	<ul style="list-style-type: none"> <li>• Most accurate method of characterizing uncertainty, as there are no inherent simplifying assumptions or activities (such as linearization)</li> </ul>	<ul style="list-style-type: none"> <li>• Very large number of samples required to characterize tails of distribution</li> <li>• Far more computationally intensive than other methods</li> </ul>
<b>Linear Mapping</b>	<ul style="list-style-type: none"> <li>• Much faster and less computationally intensive than other methods</li> <li>• Process noise provides mechanism for covariance tuning/realism adjustments</li> </ul>	<ul style="list-style-type: none"> <li>• Least accurate, especially for long propagations</li> <li>• Imposes <i>a priori</i> statistical structure upon uncertainty volume</li> <li>• Use of process noise requires careful tuning process</li> </ul>



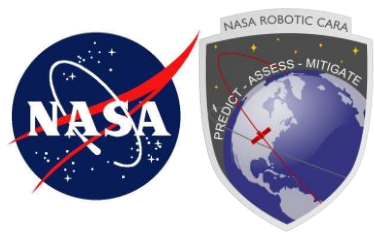
# Agenda

- **Covariance basics**
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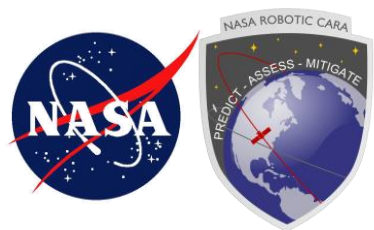
# Covariance Tuning

- **For CA, position covariance needs to be a realistic representation of the state uncertainty volume at the propagation point of interest**
- **Two aspects to this requirement**
  - Does the position error volume conform to a trivariate Gaussian distribution?
  - If so, is it of the proper dimensions and orientation?



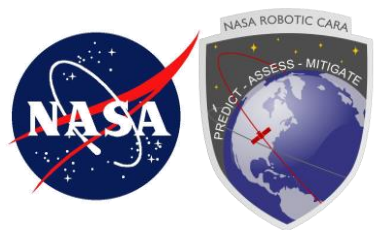
# Non-Gaussian Error Volumes

- **As propagation times increase, linearized dynamics put under more and more strain**
  - Divergence between rectilinear linearization and curvilinear orbital framework
- **Framing error volume in element space, regardless of how covariance is propagated, very strongly preserves Gaussian nature**
  - Demonstrated in Sabol (2010) paper
- **Non-Gaussian issue of significance relatively infrequently, especially for high-Pc events**
  - Shown in Ghrist and Plakalovic (2012)
- **However, does occur occasionally and can significantly alter Pc**
  - Hall *et al.* (2018)
  - Can remediate situation by performing Monte Carlo in element space
    - Monte Carlo from TCA works quite well
  - Emerging research (hope is this fall) will give criteria for when non-Gaussian issues arise with 2-D Pc



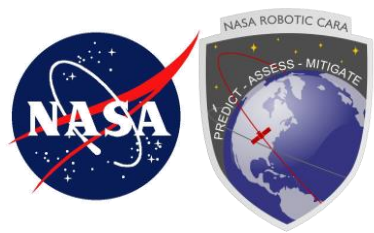
# Covariance Tuning: Covariance Realism Evaluation Method

- **Presume a precision observation available for a satellite**
  - reference orbit, or observation such as definitive solution
- **Position differences between predicted ephemeris and precision position are  $dU$ ,  $dV$ , and  $dW$** 
  - Can be collected into vector  $\epsilon$
- **Mahalanobis distance ( $\epsilon * C^{-1} * \epsilon^T$ ) represents the ratio of the difference to the covariance's prediction**
  - For a trivariate distribution, expected value is 3
- **A group of such calculations should conform to a chi-squared distribution with three degrees of freedom**
  - The method of distribution testing of groups of such calculations is used to determine if covariance properly sized



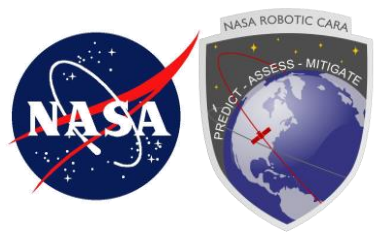
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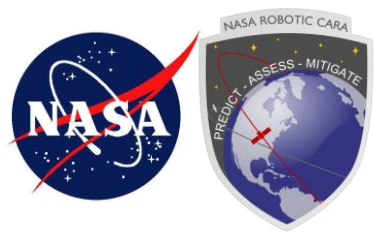
# Covariance Tuning: Covariance Irrealism Remediation

- **Three general methods for remediation**
  - Scale factor computation and application
  - Process noise definition and application
  - Consider parameter application



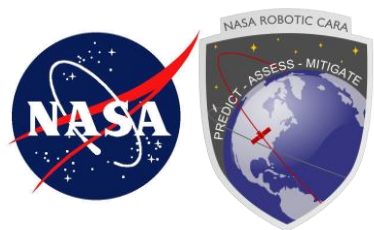
# Scale Factor Calculation

- **Used presently by JSC as part of ISS CA**
- **Mahalanobis distance calculation histories used to calculate scale factors to make covariances more realistic**
  - Can be satellite-by-satellite or for entire orbit regimes
  - Usually for a particular propagation state (e.g., 72-hour propagation)
- **Iterative (or visual) techniques used to calculate covariance scale factors that will make Mahalanobis test results match expectations**
  - Can be single scale factor by which entire matrix is pre- and post-multiplied ( $k \cdot C \cdot k$ )
  - Can be a scale factor for each of the coordinate axes ( $K \cdot C \cdot K$ , where  $K$  is a diagonal matrix of three axis-specific scale factors)
- **Factors produced by off-line analysis and then applied to each covariance in conjunctions**
  - Can be function of orbital parameters and propagation time, or simply calculated for all objects for typical propagation state of final CA decision



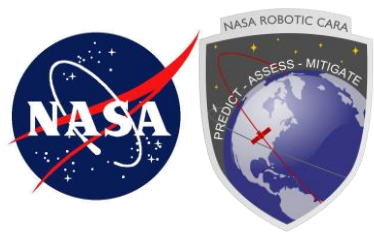
# Process Noise

- **Most commonly used with Kalman filter techniques, although can be applied to any covariance propagation**
- **Iterative approach to defining acceleration variances**
  - Originally produced as technique for compensating for inadequate geopotential solution when processing restrictions limited geopotential fidelity
  - Now employed more as a general method to compensate for covariance inadequacies
- **Variances used to define Q matrix, which is propagated through process noise state transition matrix  $\Gamma$** 
  - Added to propagated covariance
- **If done correctly, construct should work for all propagation states**
  - Advantage over scale factor approach, for which separate scale factor needed for each (significantly) different propagation state



# ASW Covariance Propagation and Consider Parameters

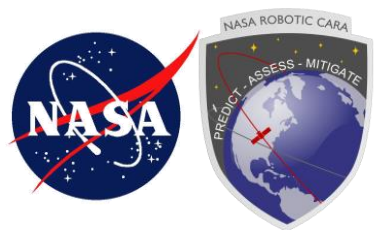
- **Covariance in VCM is unaltered covariance**
- **When propagating VCM covariance, the propagator**
  - Scales the covariance by the weighted RMS if it is greater than unity
    - $C^* = C * WRMS^2$
    - Was an early attempt to improve covariance realism; not clear this is still a good idea
  - Applies the consider parameter to the ballistic coefficient variance
    - $C^*(7,7) = C^*(7,7) + Cpd^2$
    - More later on how this value is determined
  - May apply a consider parameter to the solar radiation pressure variance
    - $C^*(8,8) = C^*(8,8) + Cps^2$  (would be position (9,9) on VCM)
    - Presently not used (Cps set to 0)
  - Propagates the altered covariance using linearized dynamics
    - $\Phi * C^{**} * \Phi^T$
  - Converts propagated matrix from equinoctial to Cartesian coordinates



# Altered Covariance Positions

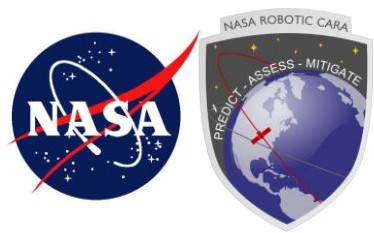
- **Ballistic coefficient consider parameter (DCP) applied to ballistic coefficient variance (orange)**
- **If used, solar radiation pressure consider parameter applied to solar radiation pressure variance (purple)**

	U (m)	V (m)	W (m)	Udot (m/s)	Vdot (m/s)	Wdot (m/s)	B (m <sup>2</sup> /kg)	AGOM (m <sup>2</sup> /kg)
U	6.84E+01	-2.73E+02	6.38E+00	2.76E-01	-7.14E-02	8.75E-03	-3.83E-02	-3.83E-02
V	-2.73E+02	1.10E+05	3.23E+01	-1.17E+02	-8.99E-02	2.51E-02	-1.28E-01	-1.28E-01
W	6.38E+00	3.23E+01	4.47E+00	-3.26E-02	-6.83E-03	1.81E-03	-3.73E-03	-3.73E-03
Udot	2.76E-01	-1.17E+02	-3.26E-02	1.24E-01	1.10E-04	-2.47E-05	1.46E-04	1.46E-04
Vdot	-7.14E-02	-8.99E-02	-6.83E-03	1.10E-04	7.57E-05	-9.39E-06	4.10E-05	4.10E-05
Wdot	8.75E-03	2.51E-02	1.81E-03	-2.47E-05	-9.39E-06	2.06E-05	-4.39E-06	-4.39E-06
B	-5.07E-03	1.30E+00	4.34E-05	-1.38E-03	7.97E-07	7.26E-07	1.64E-05	-6.28E-07
AGOM	-3.83E-02	-1.28E-01	-3.73E-03	1.46E-04	4.10E-05	-4.39E-06	-6.28E-07	2.31E-05



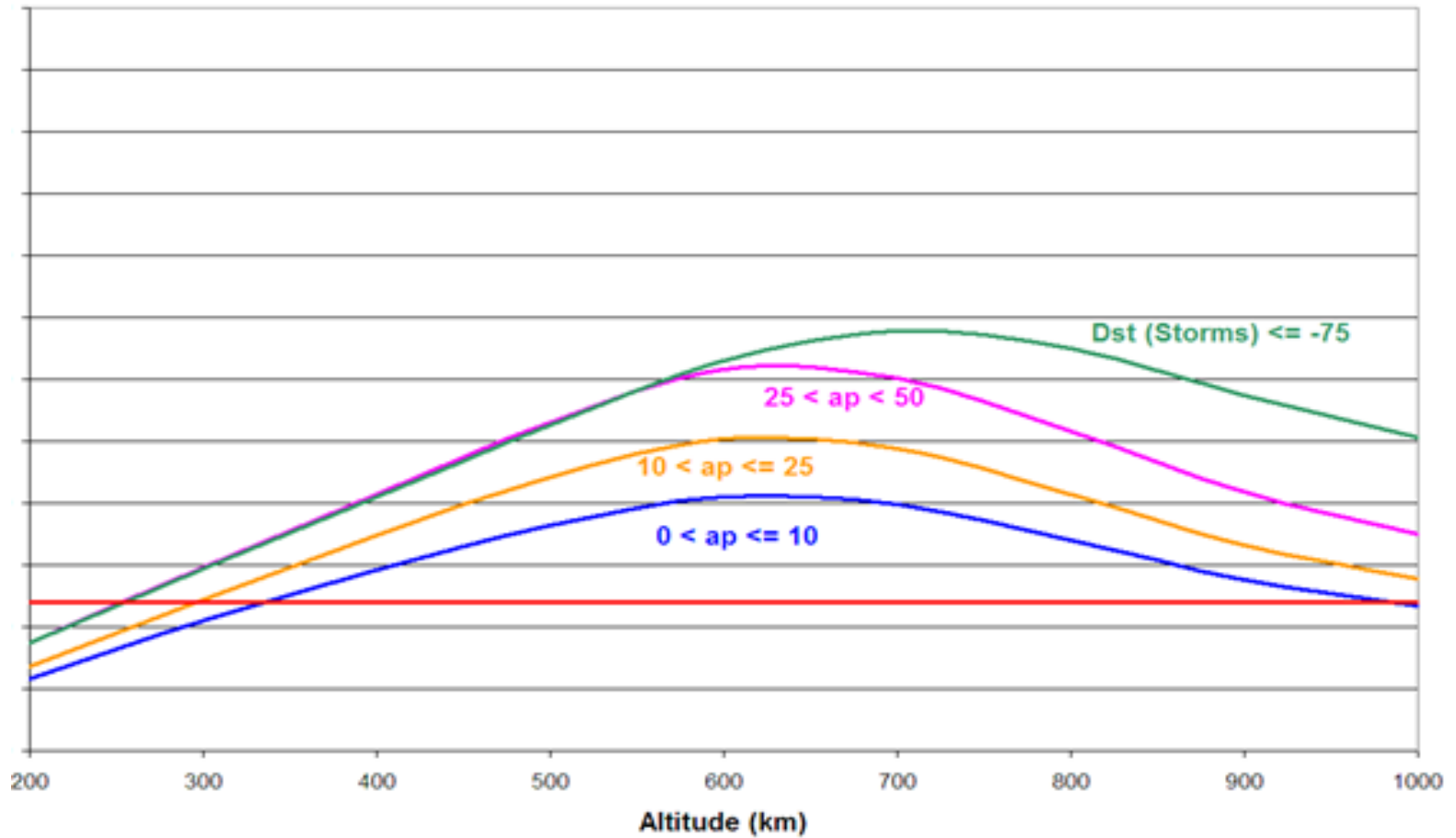
# Dynamic Consider Parameter (DCP)

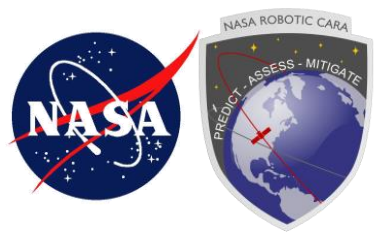
- **Specifies global error in the atmospheric density forecast**
  - Parameterizes percent RMS error in terms of
    - Satellite height (perigee altitude)
    - Geomagnetic activity ( $a_p$  and Dst)
  - Density forecast error combination of solar/geomagnetic indices and DCA
    - Directly compared numerous forecast densities to actual density
    - Discretized heights from 200 km to 1000 km averaged over lat / lon
  - Sampled high, low, and medium solar cycles (2001, 2005, 2013)
    - Found most variation parameterizable via  $a_p$  conditions (versus  $F_{10}$ )
    - Functions optimized for 3-day predictions—this is the tuning point!
- **Determines satellite-specific frontal area variation in prediction**
  - Quantifies ballistic coefficient RMS error through satellite histories
    - Looks back in time up to a year in most cases (upfront preprocessing)
    - Ascertains error / target for 3-day predictions (accounting for time-lags)
- **Combine the two uncertainty components to obtain DCP value**
  - Additive in variance sense as the root sum of squares



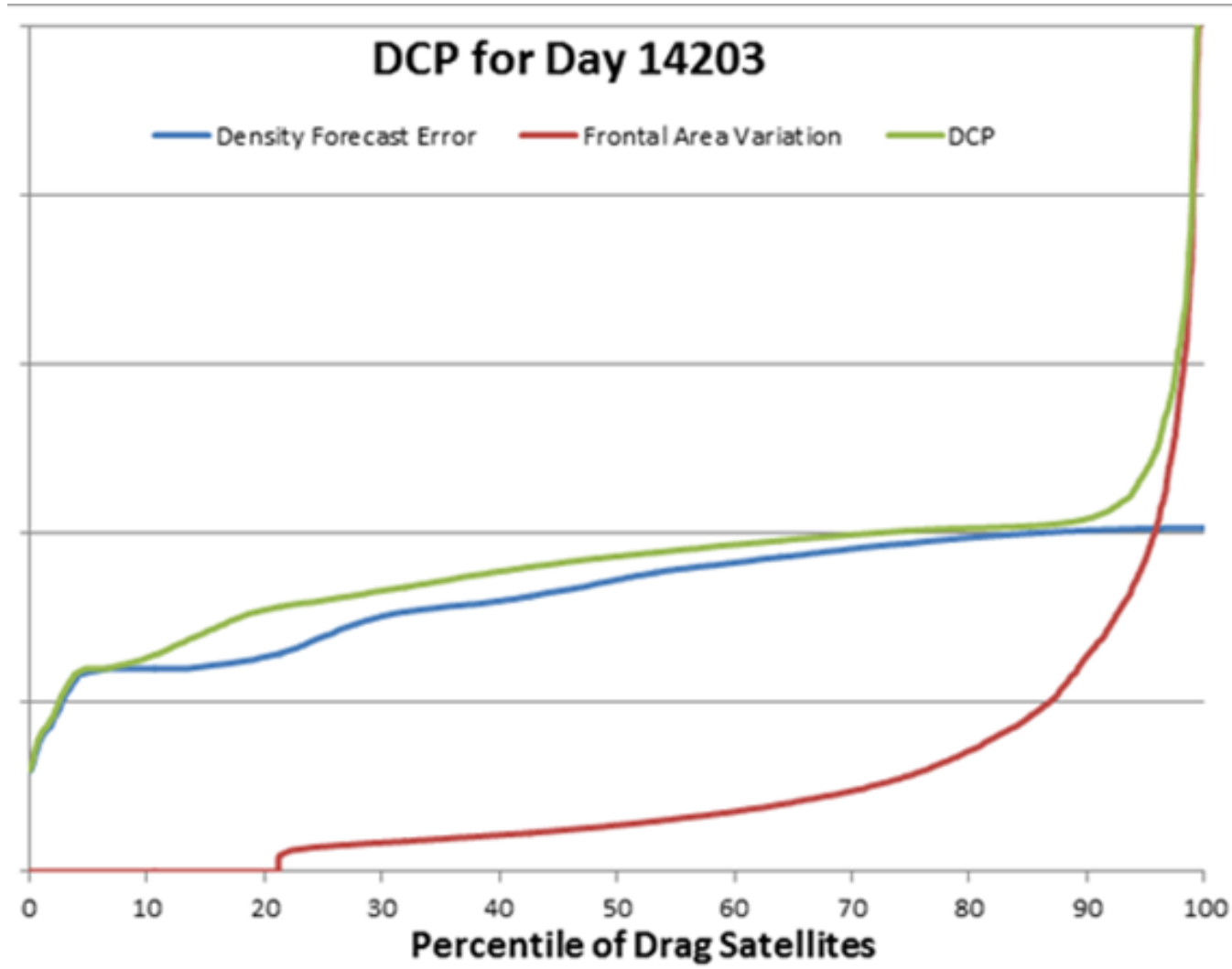
# B Consider Parameter Values

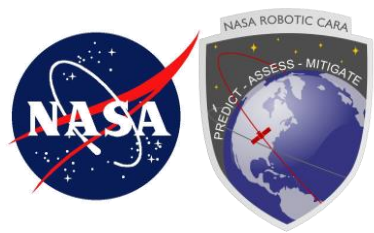
dRho STD Predict ap & Dst





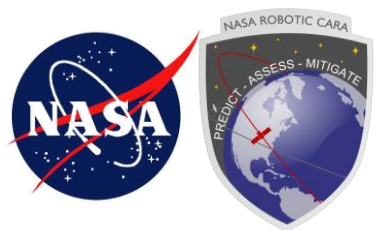
# DCP Components RMS Uncertainty





# Agenda

- **Covariance basics**
- **Use of covariance in probability of collision ( $P_c$ ) calculation**
- **Covariance generation and propagation methods**
- **Covariance tuning**
- **Covariance realism improvement methods**
- ➔ • **Covariance theory compatibility**



# Covariance Theory Compatibility

- **Batch covariance is governed by the amount and quality of tracking data used in the OD**
- **Propagated covariance is a product of the particular propagation technique used and tuning applied**
  - Tuning itself a function of the adequacy of the OD force modeling
- **Thus, important that covariance be generated from same OD basis that produced the state estimate**
- **This is not possible for O/O ephemerides that lack a covariance**
  - Forced to use O/O state estimate and ASW covariance (or a synthesized covariance when no ASW covariance exists)
  - Such a covariance a questionable representation of O/O ephemeris error
    - $\sigma_{O/O}^2 = \sigma_{ASW}^2 + \sigma_{Diff}^2$
    - The difference variance is unknown, so using an ASW covariance with an O/O ephemeris understates the uncertainty but by an unknown amount