



James Webb Space Telescope Optical Telescope Element / Integrated Science Instrument Module (OTIS) Cryogenic Vacuum Test

Part I: Thermal Architecture

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Introduction to the JWST OTIS CV Test Lecture Series



- This lecture series discusses the cryogenic vacuum (CV) testing of the James Webb Space Telescope (JWST) Optical Telescope Element / Integrated Science Instrument Module (OTIS)^[1,2] from July – October, 2017 at NASA Johnson Space Center

- There are four parts to this series:
 - Part I: Thermal Architecture
 - Part II: Thermal Analysis
 - Part III: Preparations for Off-Nominal Events
 - Part IV: Lessons Learned

- Objectives of this lecture series:
 - Familiarize the audience with the James Webb Space Telescope architecture
 - Understand the thermal challenges to executing the most complex cryogenic vacuum test ever undertaken by NASA
 - Act as a guideline for planning future system-level thermal vacuum tests for large cryogenic missions



The James Webb Space Telescope: Introduction



- Upon launch in 2021, the James Webb Space Telescope will become the world's most powerful general-purpose space observatory
 - Scientific successor to the Hubble Space Telescope
 - Optimized to observe in near-to-mid infrared wavelengths (0.6 – 28 μm)



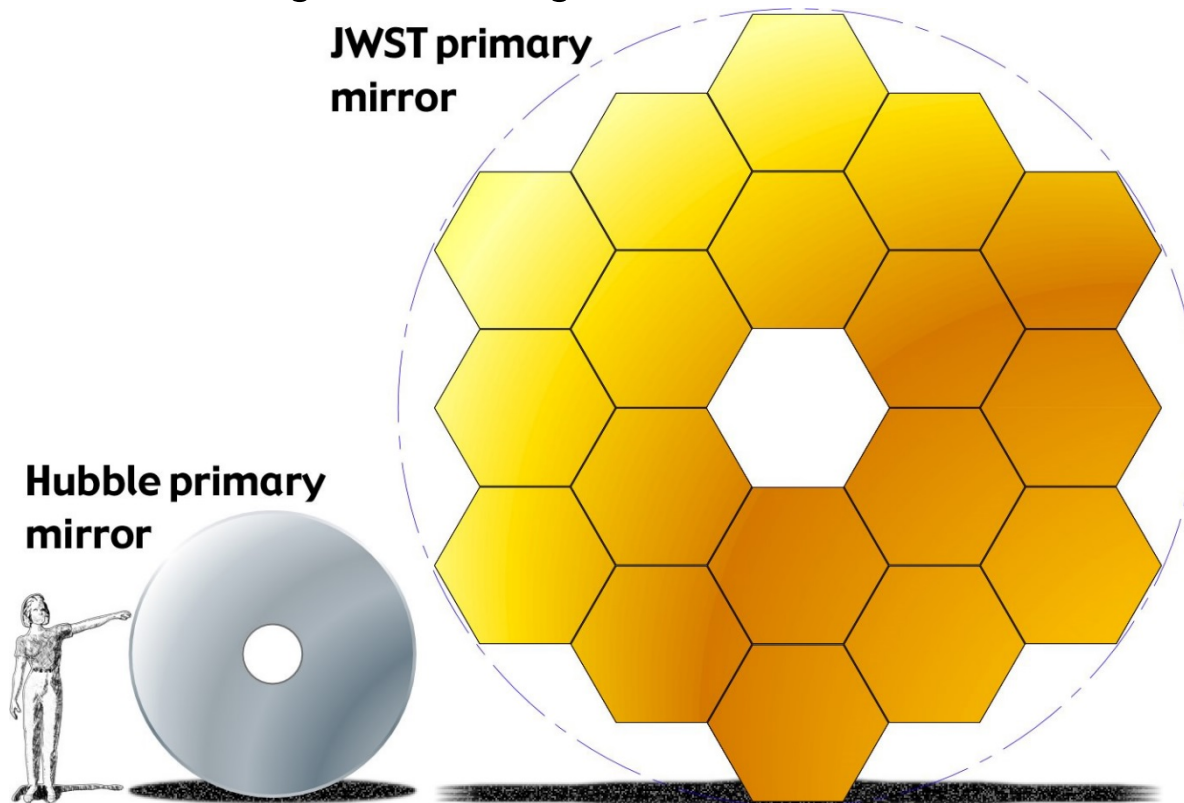
JWST Full-scale model at SXSW in Austin, TX, 2013. Image source: NASA/JWST



JWST vs. Hubble Primary Mirror Comparison



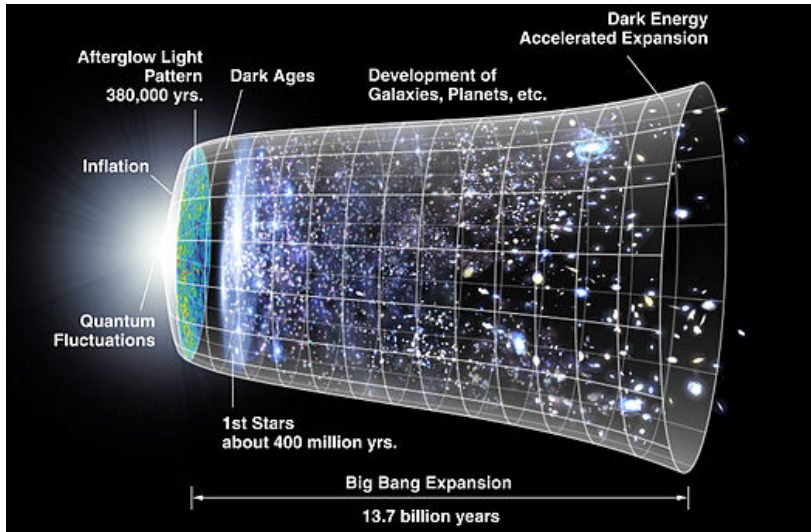
- JWST's primary mirror is 6.5 m in diameter and has 25 m² of light-collecting area
 - This is powerful enough to see the heat of a bumblebee on the moon from Earth!
- By comparison, the Hubble Space Telescope's primary mirror is 2.4 m in diameter with 4.5 m² of light-collecting area



Source: NASA/JWST

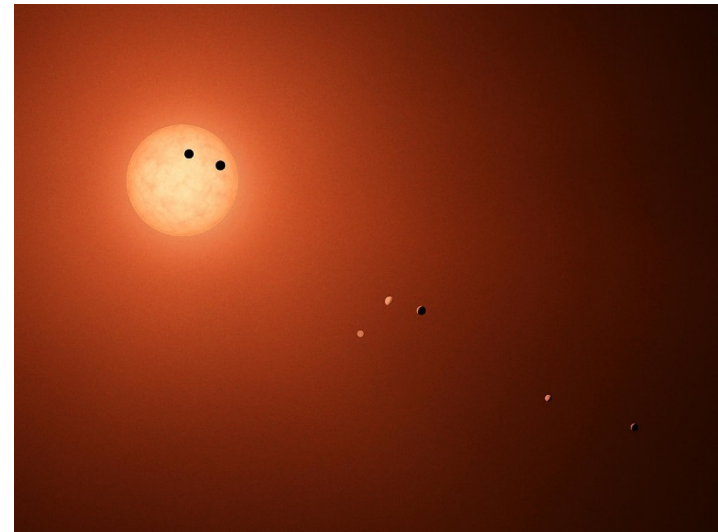


JWST Science [3]



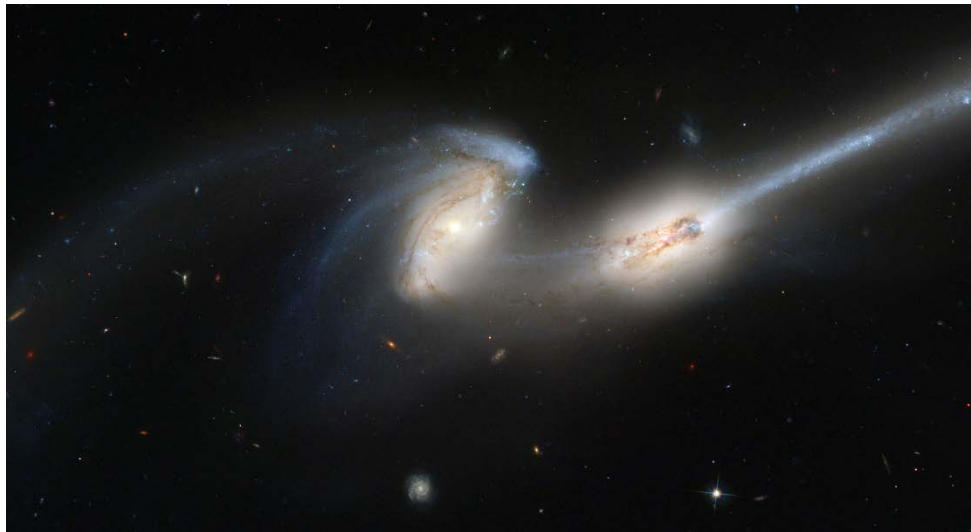
First Light and Reionization

Source: NASA/WMAP



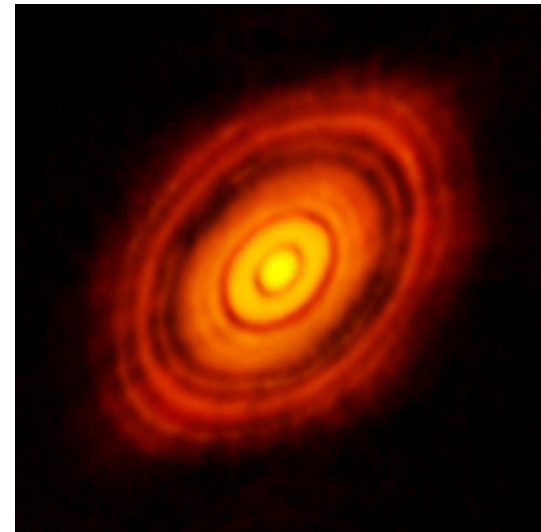
Planetary Systems and the Origins of Life

Source: NASA/JPL-Caltech



Assembly of Galaxies

Source: NASA/ESA

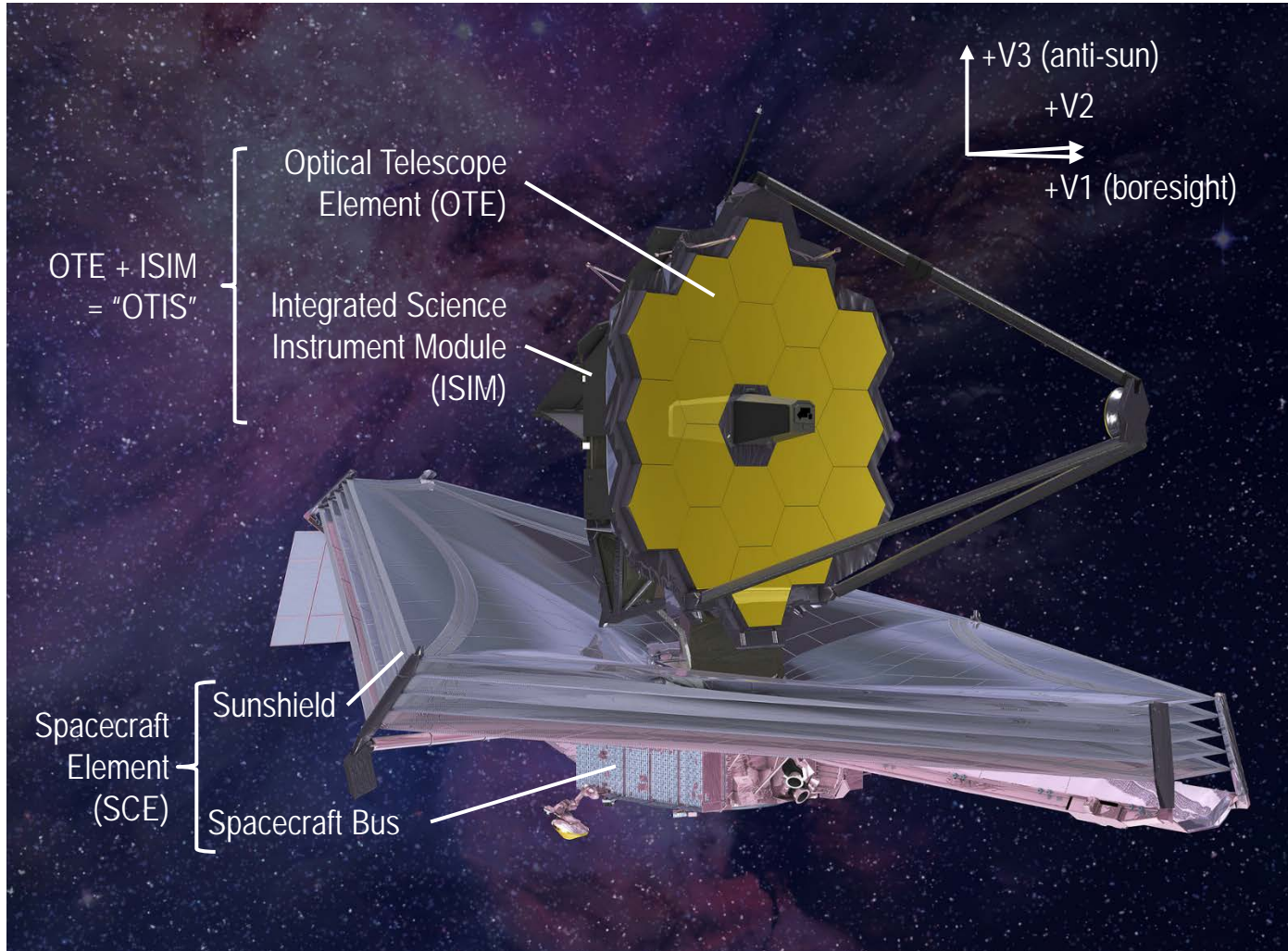


Birth of Stars and Protoplanetary Systems

Source: ALMA Observatory



Major System-Level Assemblies of the James Webb Space Telescope



Source: NASA/JWST



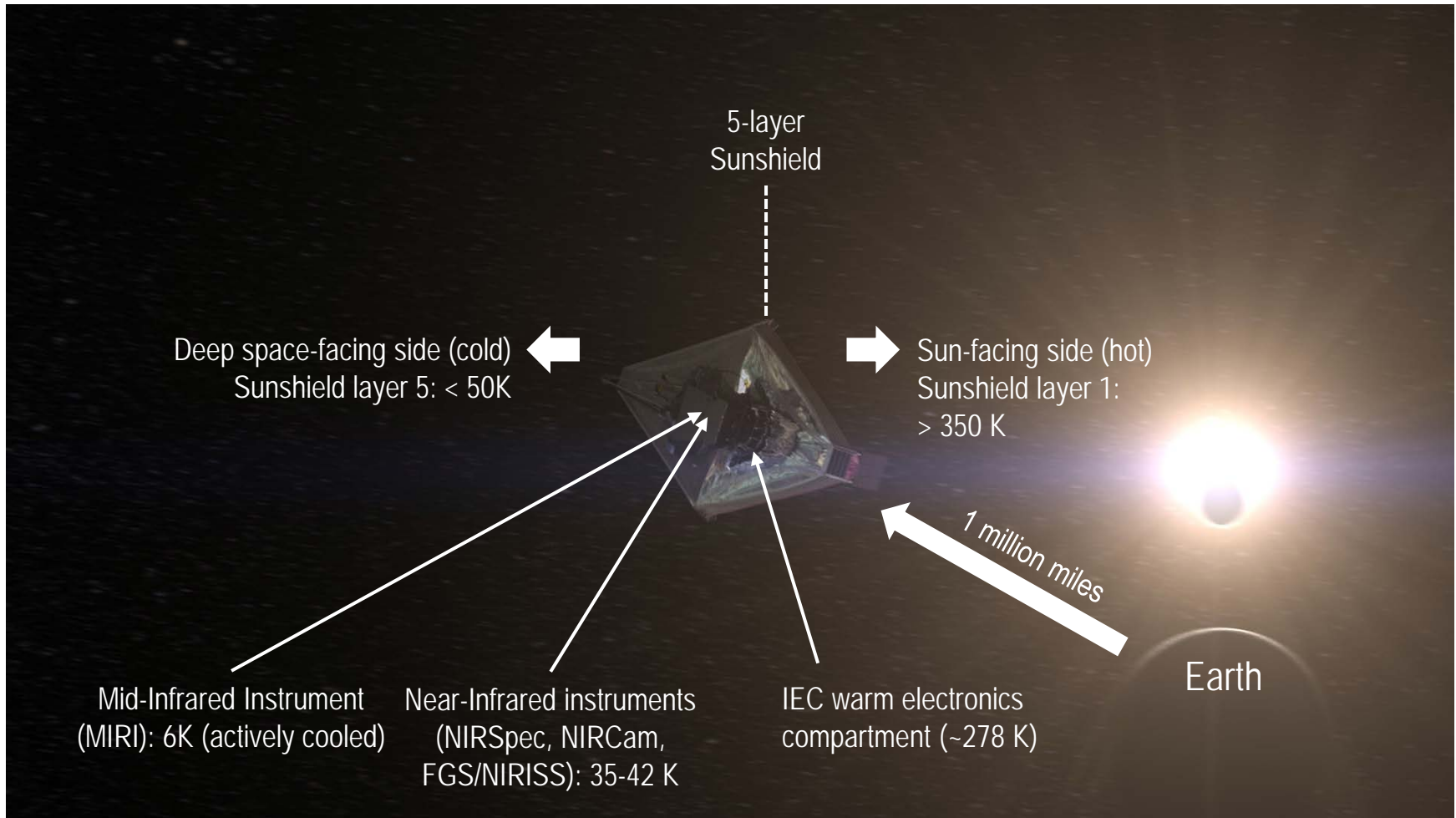
Completed "OTIS" Element at NASA Goddard Space Flight Center SSDIF Clean Room



Source: NASA/JWST



Thermal Environment in Flight



Source: svs.gsfc.nasa.gov



OTIS CV Thermal Test Objectives [4]



- Preserve hardware integrity upon transition to cryogenic thermal balance (cryo-balance) conditions and transition back to ambient temperatures by respecting all imposed limits and constraints (L&Cs)
- Achieve the simulated on-orbit payload temperature levels and stability for optical, mechanical, and instrument tests
- Predict and measure thermal balance test data for model crosscheck, both on ISIM and OTE components
- Achieve a workmanship thermal conductance assessment of the flight instrument heat straps which for the first time would be connecting all the payload flight instruments and radiators
- Achieve test timeline optimization by executing the OTIS CV cooldown and warmup in a time-efficient manner



What's the Importance of Thermal Vacuum Testing?

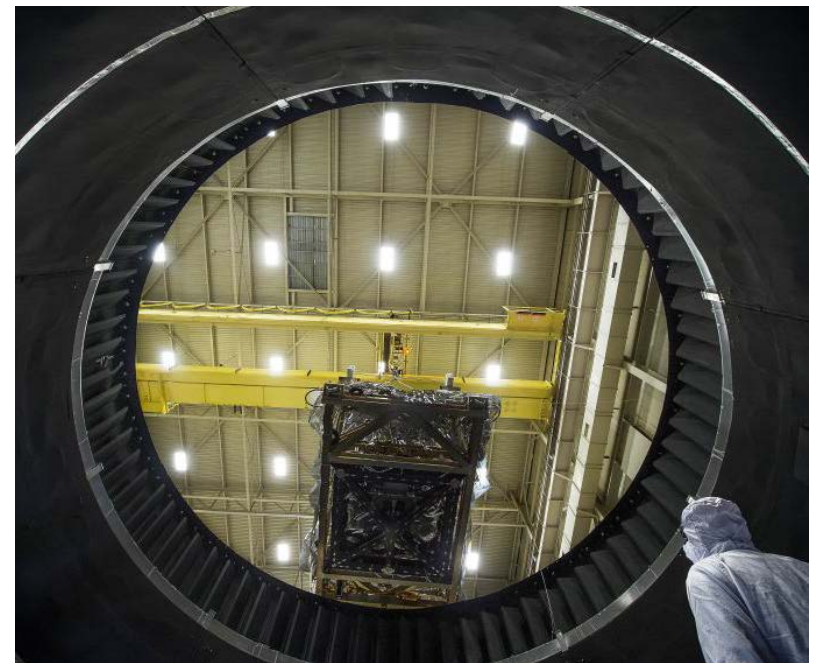


Thermal testing is done in a vacuum chamber at *margined* temperature extremes and is designed to verify workmanship, demonstrate performance, and collect data to be used in correlating thermal models

Two types of testing are performed:

Thermal balance plateaus: thermal environment is set, and spacecraft must achieve energy balance with environment. Balance criteria met from achieving temp. rate-of-change requirement on components. Thermal data collected is used to verify predictive accuracy of thermal models

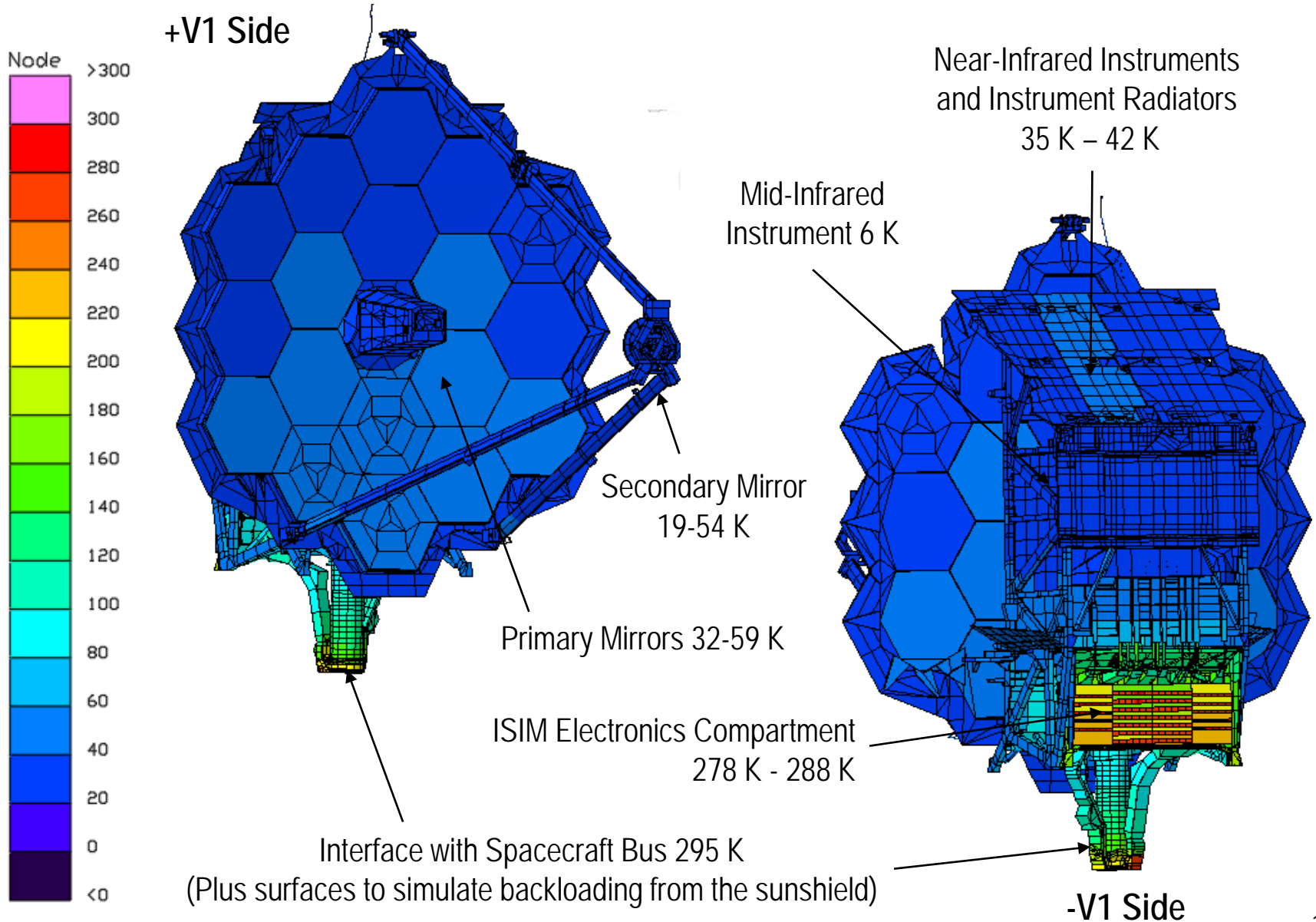
Thermal vacuum cycles: Quality assurance test to take hardware beyond its operational temperatures and ensure it will survive temperature extremes: used to verify workmanship on components



*Cryo-vacuum testing of ISIM at NASA Goddard Space Flight Center's Space Environment Simulator Chamber
Source: NASA/Chris Gunn*

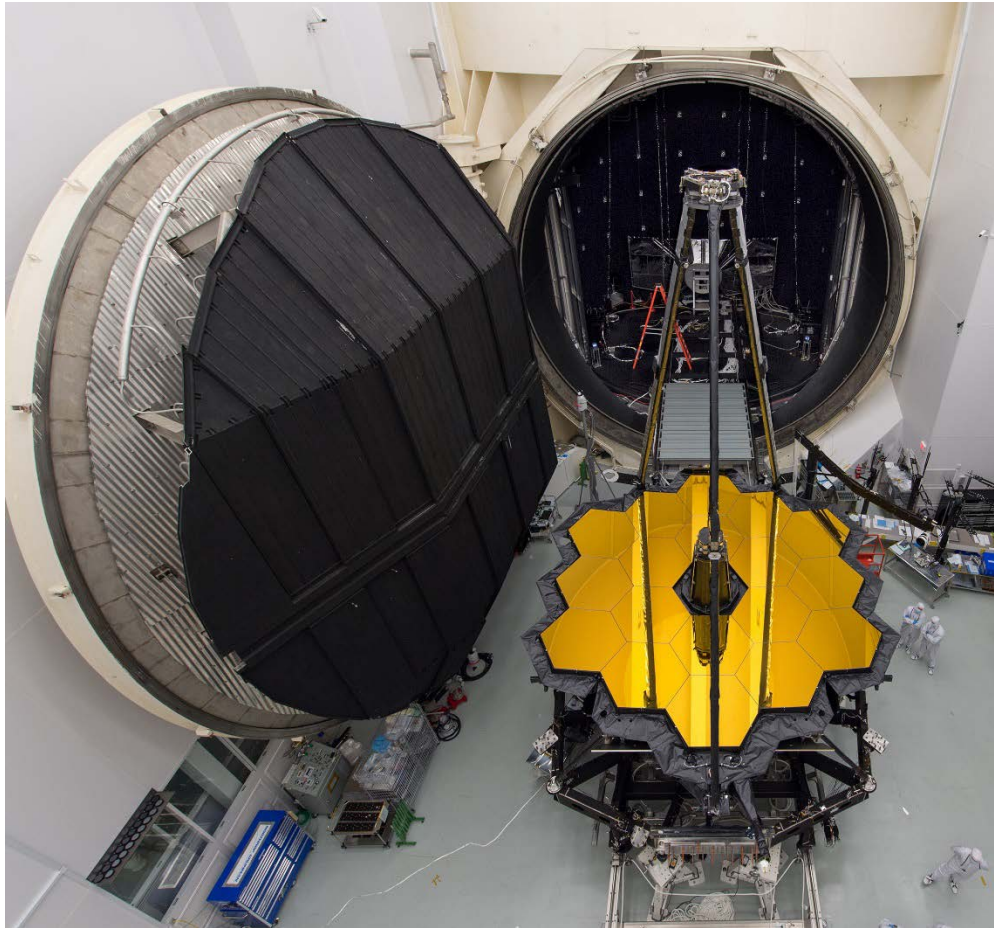


What Are Our Temperature Goals on OTIS?





How Do We Replicate JWST's Flight Thermal Environment in Test ?

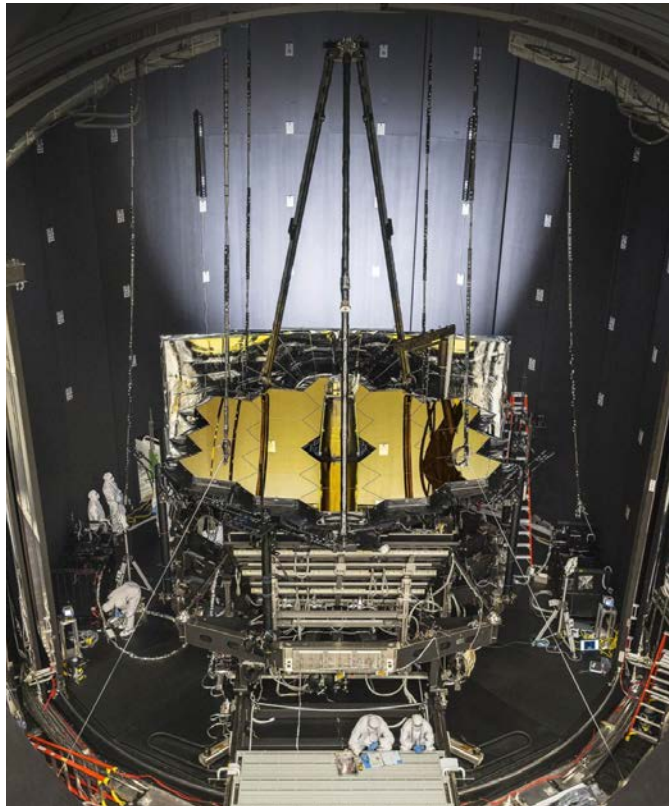


Source: NASA/JWST

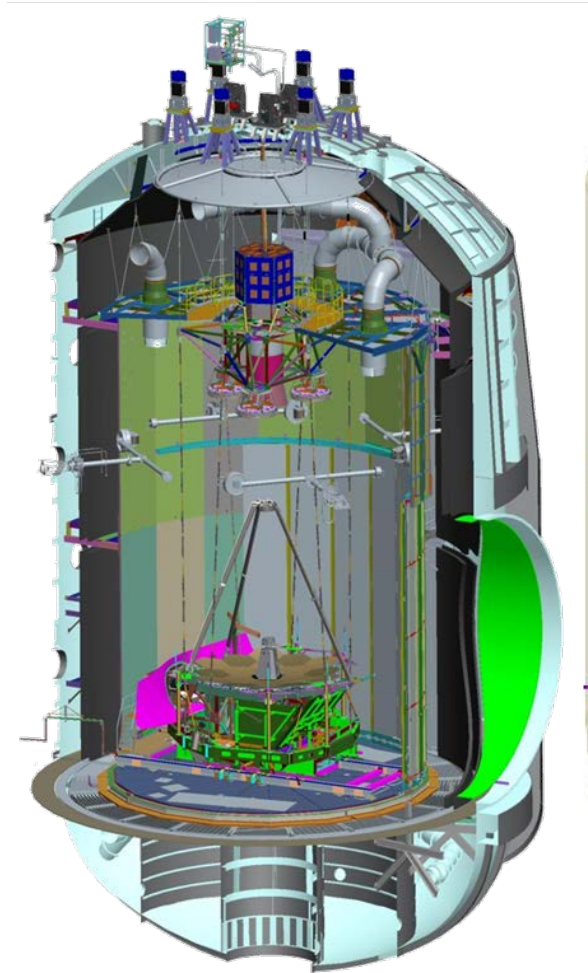
- Use one of the largest thermal vacuum chambers in the world (NASA Johnson Space Center's Chamber A)
 - Unfortunately, even this chamber is not large enough to fit all of JWST, so we need to test in separate system-level assemblies (OTIS being the major cryogenic test)
- Install a gaseous helium shroud to lower the payload temperatures to 20K, and an LN2 shroud to lower the overall environmental loads on the helium shroud/refrigerator ^[5]
- Install GSE to simulate heat from the flight spacecraft bus



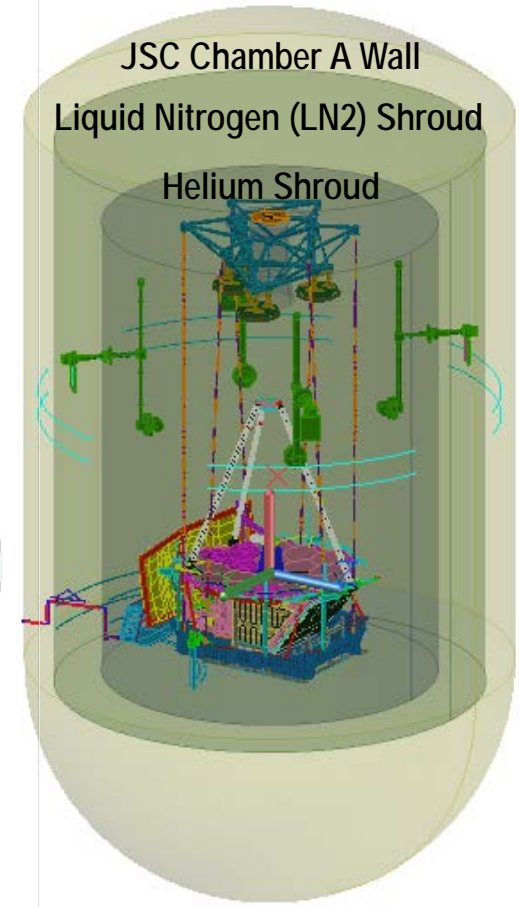
OTIS CV Test Setup Inside Chamber A: Three Different System-Level Representations



Physical Hardware



CAD Model

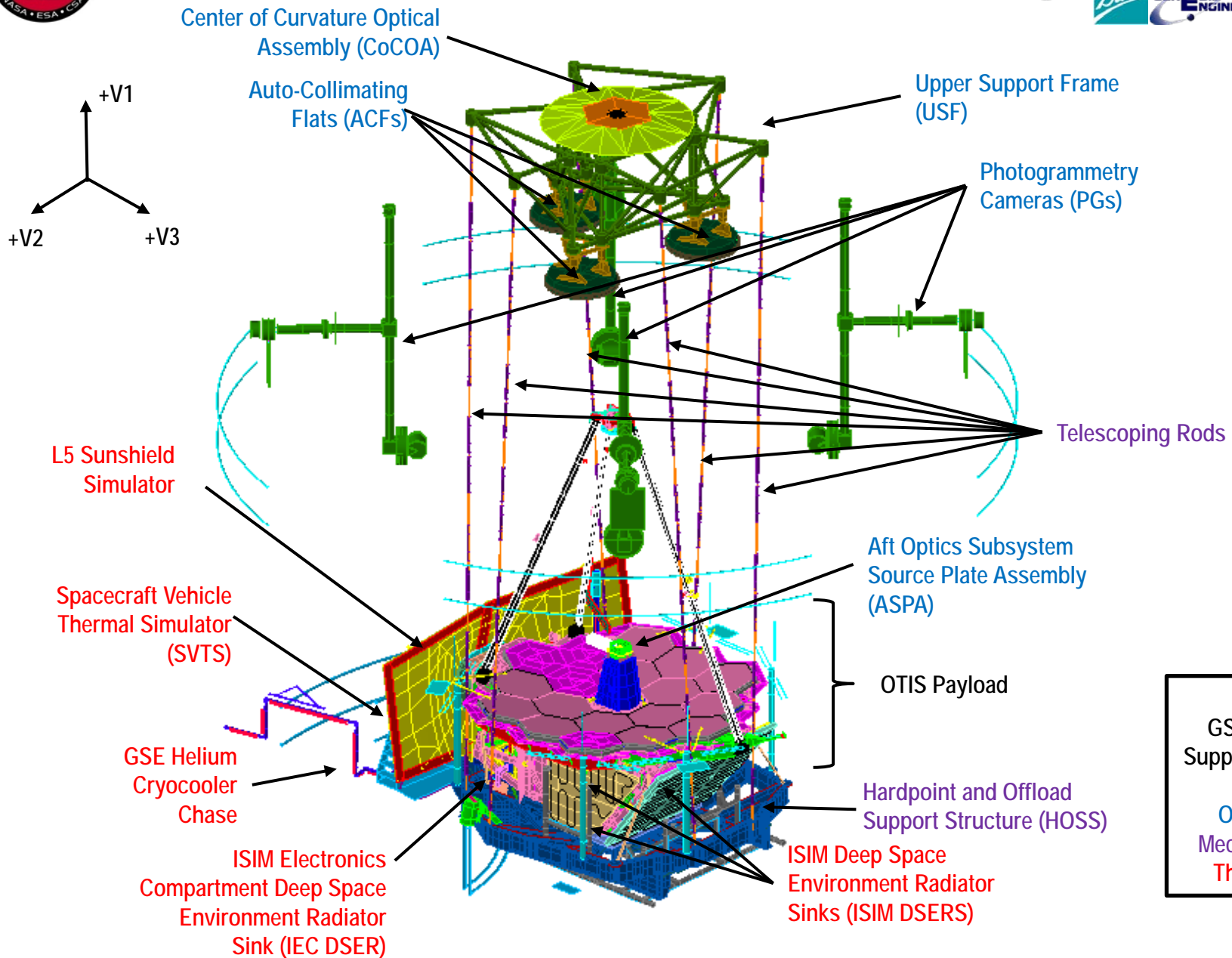


Thermal Model

Image sources: NASA/JWST



JWST OTIS CV Test Setup: Inside the Helium Shroud



Source: NASA/JWST

LEGEND

GSE = Ground Support Equipment

Optical GSE

Mechanical GSE

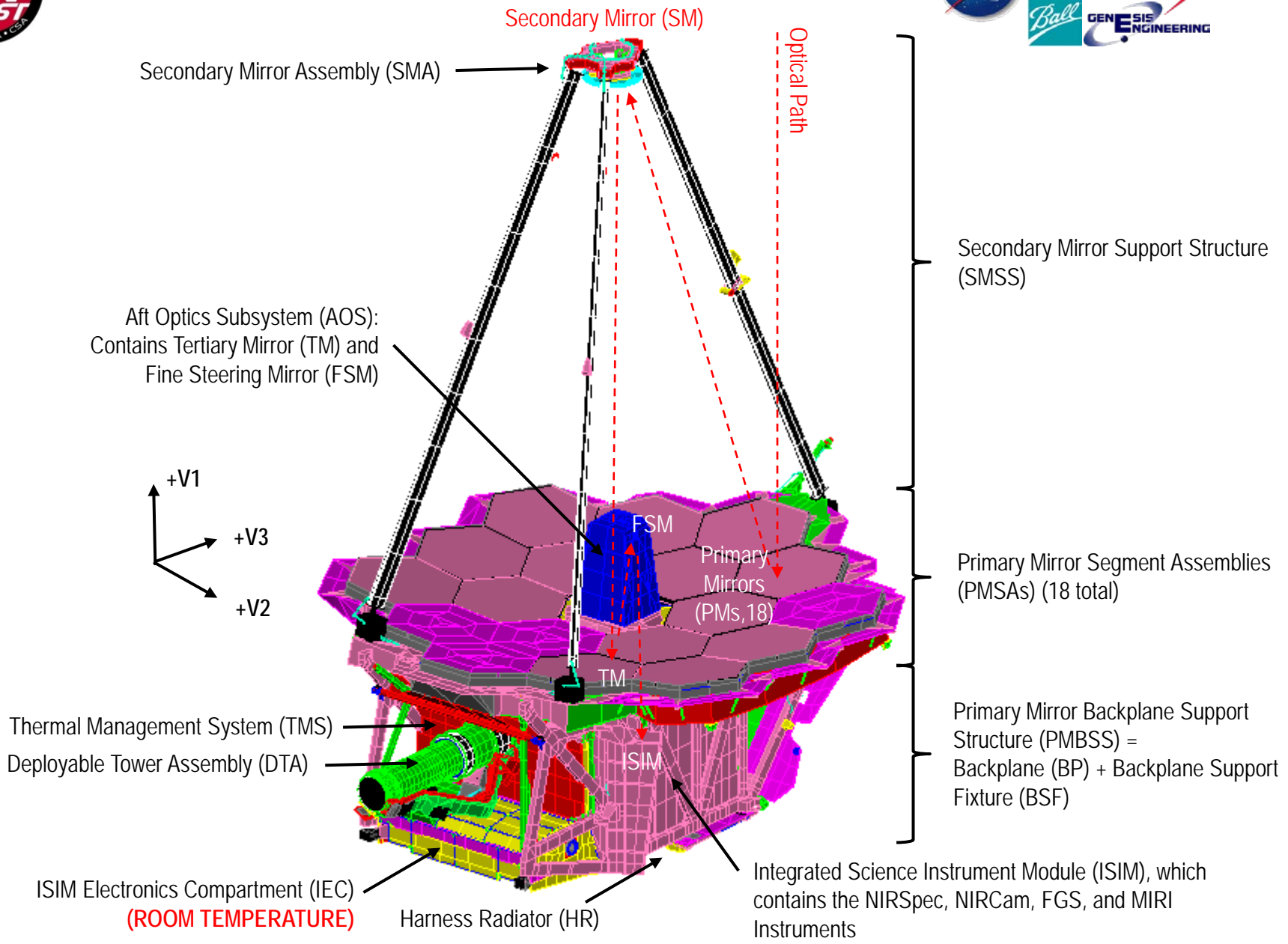
Thermal GSE



JWST OTIS CV Test Setup: Payload Configuration



Source: NASA/JWST





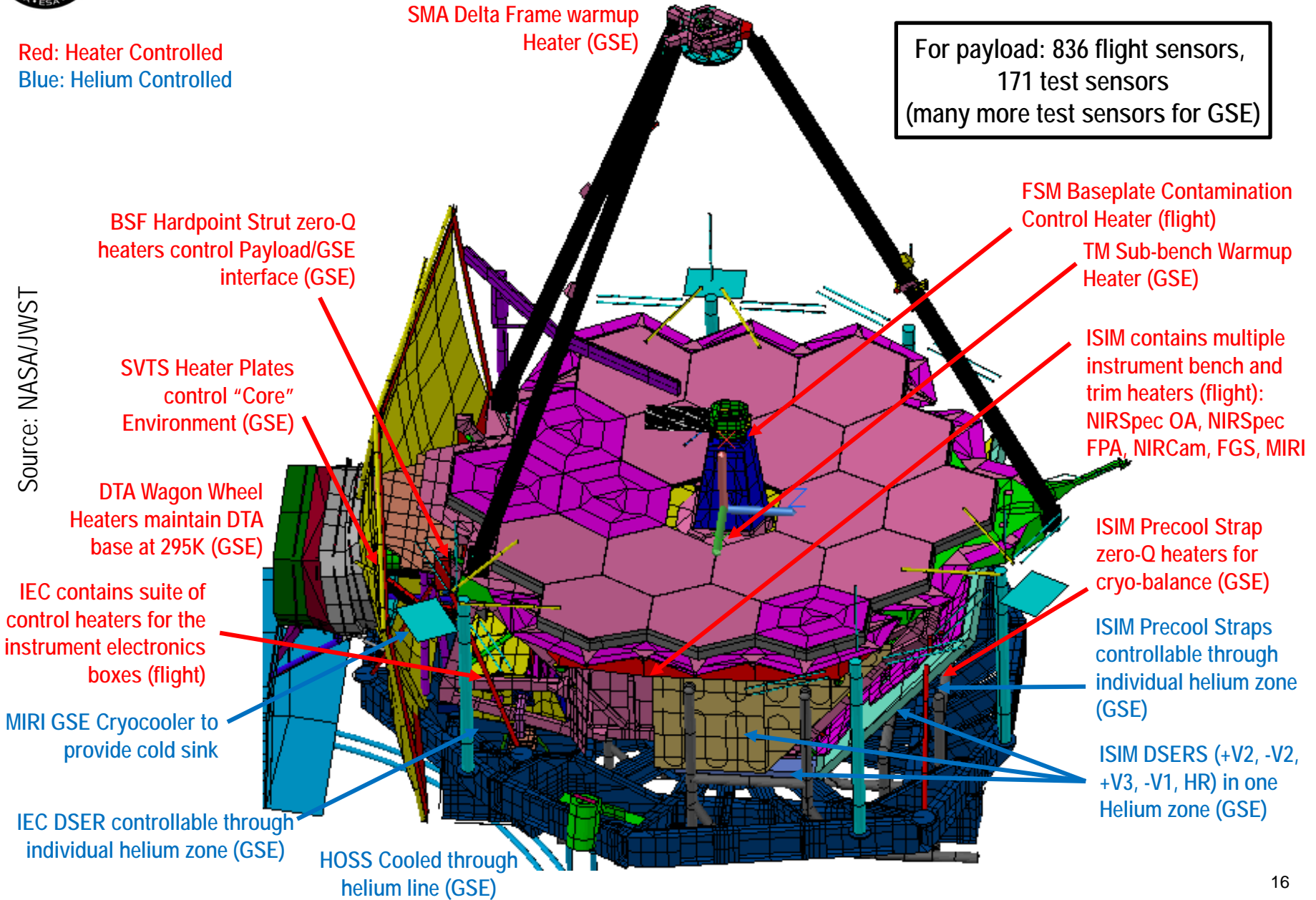
OTIS Thermal Control Hardware [6]



Red: Heater Controlled
Blue: Helium Controlled

For payload: 836 flight sensors,
171 test sensors
(many more test sensors for GSE)

Source: NASA/JWST

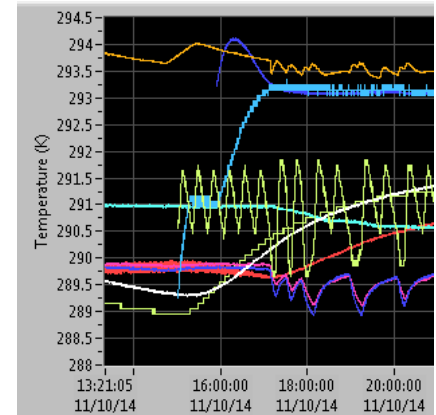




GSE Considerations for the Cryogenic Test Environment [7]



- A robust thermal instrumentation plan was developed with multiple systems to rigorously interpret cryogenic test results
 - Calibrated diodes, precise data acquisition units for accuracy/resolution through range of test temperatures
 - Radiometers to measure localized heat sources
 - Calorimeters for understanding radiative boundaries and icing
- Thermal balance test required precise control of boundary heat leaks on the mW scale, and optical / instrument tests required management of stray light entering optical path
 - Stationary penetrations on Helium shroud closed out with single layer insulation (SLI) or multi-layer insulation (MLI)
 - Specialized systems of light-tight baffles, shell structures, and MLI used for shroud penetrations which moved due to cryo-shift (e.g. Down / Telescoping Rods) or mechanism operations (e.g. photogrammetry cameras)
 - Harnessing from external environment was anchored to increasingly colder thermal sinks to reduce stray light into chamber



Screenshot of Harris TTS Data Acquisition System
Source: Harris Corp.



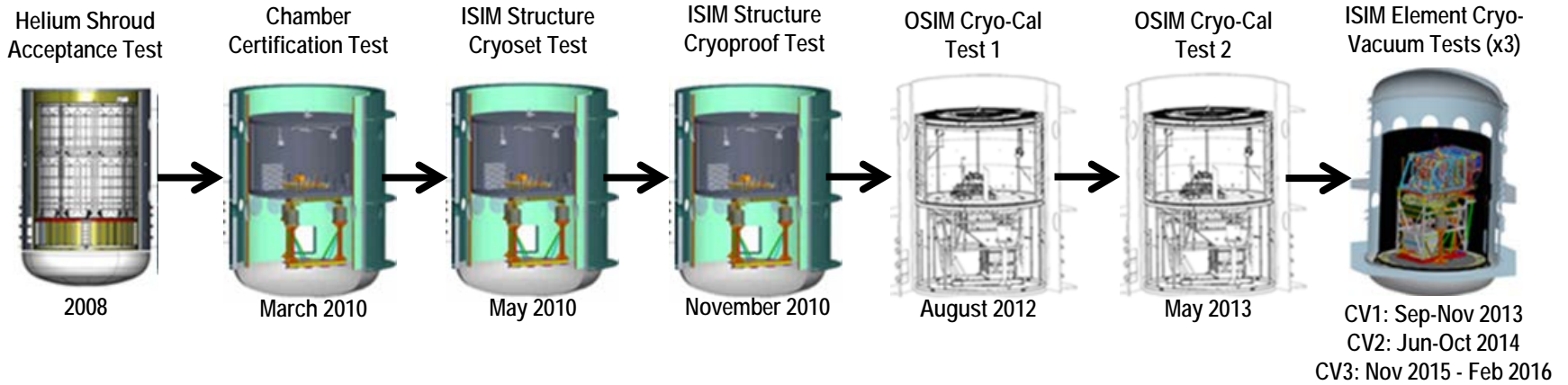
Closeout of Down Rods
Source: Harris Corp.



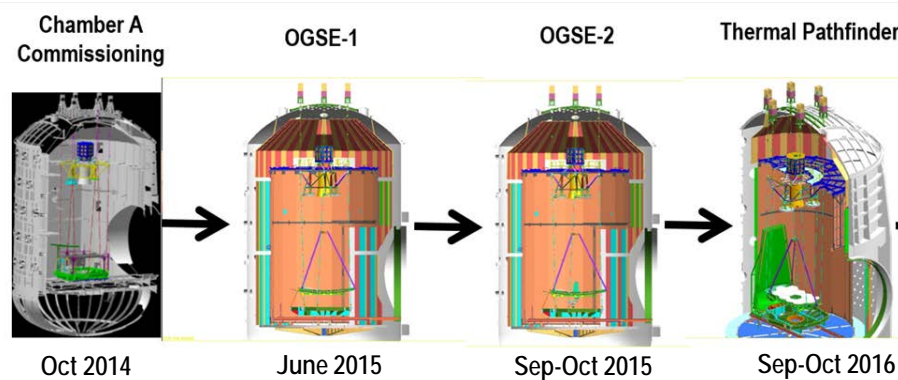
How Did We Prepare for OTIS?



Major ISIM Element Thermal Vacuum/Thermal Balance Tests (SES Chamber, NASA GSFC) [8]

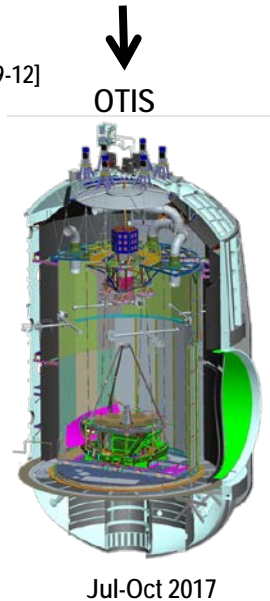


Major OTE Thermal Vacuum/Thermal Balance Tests (Chamber A, NASA JSC) [9-12]



- OTIS Analytical Models:**
- Contamination
 - Cryocooler
 - Mechanical / Dynamics
 - Optical / Stray Light
 - Spacecraft Sim / Software
 - Thermal
 - Thermal Distortion

Multi-year Development / Iterative Process





What Are Our Requirements?



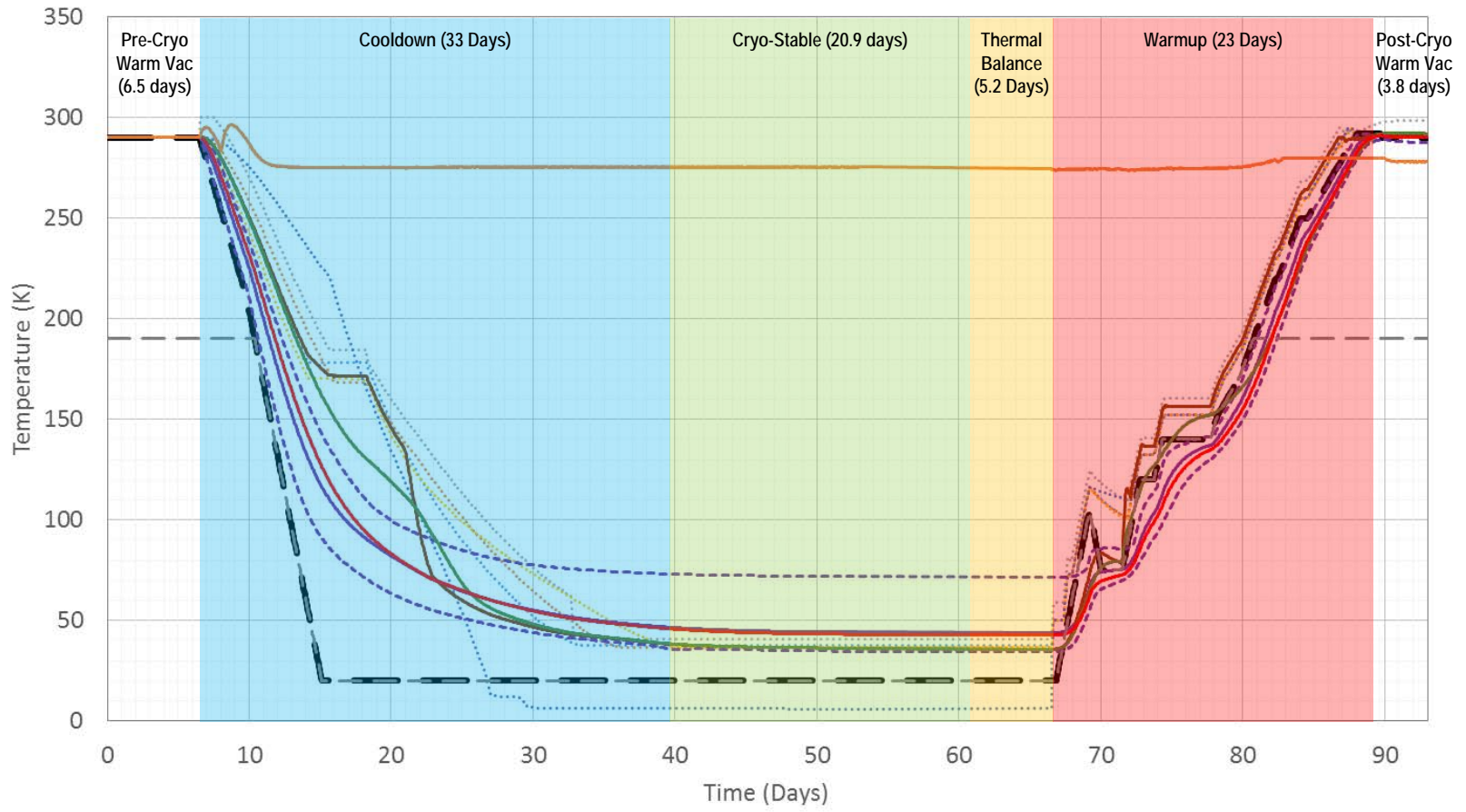
Constraints are put in place to avoid actions, conditions, or events, which if realized, will result in damage to flight hardware.

Limitations are put in place to avoid actions, conditions, or events, which have the potential for temporarily impacting performance or resulting in loss of test time.

- **For the Thermal Subsystem, there were 84 constraints and 8 limitations out of more than 1,000 total for the OTIS test**
 - Most thermal constraints and limitations were designed to avoid contamination, overstressing of structural elements and instruments. They defined absolute temperature limits, rates of change, gradients within structures, instruments, and temperature relationships between instruments, optics, thermal boundaries, usage of heaters
- **The OTIS thermal team installed alarms to monitor and prevent any exceedances of L&Cs**
 - Separate monitoring systems were used for flight and GSE sensors
 - FUSION, an in-house system developed at NASA GSFC, was employed to visualize both flight and GSE sensor data as it pertained to thermal-specific L&Cs
 - An alarm limit philosophy was developed to provide margin and time to respond on components which had L&Cs levied against them, but which did have sensors to directly measure their temperature against L&Cs



Full OTIS Predicted Profile



- NIRCam Bench
- NIRSpec OA
- NIRSpec FPA
- FGS/NIRISS Bench
- MIRI Bench
- Helium Shroud/ISIM DSER Average
- PMBSS Structure Max
- PMBSS Structure Avg
- PMBSS Structure Min
- FSM Substrate
- TM Substrate
- Primary Mirrors Avg
- IEC Equipment Panel Average
- IEC DSER Average

Source: NASA/JWST



Part I Summary



- In this lecture, we reviewed the mission of James Webb Space Telescope's OTIS element and its thermal architecture
 - Science objectives for OTIS
 - Constituent components (flight and GSE)
 - Justification for and method of thermal vacuum testing
 - Thermal test objectives and requirements

- In the next lecture, we will focus on the development of the thermal test methodology via thermal analysis
 - Assembly of OTIS system-level model
 - Test profile generation via limit and constraint "feedback loop"
 - Accommodations for contamination and structural concerns



Reference: Acronyms (Page 1)



Acronym	Definition	Acronym	Definition
AOS	Aft Optical System	ESA	European Space Agency
ACF	Auto-Collimating Flat	FGS	Fine Guidance Sensor
ADIR	Aft Deployable ISIM Radiator	FIR	Fixed ISIM Radiator
ASPA	Aft Optical System Source Plate Assembly	FPA	Focal Plane Arrays
BP	Back Plane	FSM	Fine Steering Mirror
BSF	Backplane Support Fixture	GSE	Ground Support Equipment
CoCOA	Center of Curvature Optical Assembly	GSFC	NASA Goddard Space Flight Center
CPP	Cryo-Pumping Panels, cold panels between the Helium and LN2 shrouds at NASA JSC	HOSS	Hardpoint and Offload Support Structure
CSA	Canadian Space Agency	IEC	ISIM Electronics Compartment
CTE	Coefficient of thermal expansion	IR	Infrared
CV	Cryogenic Vacuum	ISIM	Integrated Science Instrument Module, which contains the Science Instruments (SIs)
$\Delta T, \Delta t$	Change in temperature; change in time	JSC	NASA Johnson Space Center
DTA	Deployable Tower Assembly	JWST	James Webb Space Telescope
DSERS	Deep Space Environment Radiative Sink	K	Kelvin
EC	European Consortium	L&Cs	Limits and Constraints



Reference: Acronyms (Page 2)



Acronym	Definition	Acronym	Definition
L5	Layer 5 Sunshield simulator	POM	Instrument Pick-Off Mirror
LN2, N2	Liquid Nitrogen; Gaseous Nitrogen	PM	Primary Mirror(s)
LRM	Launch Release Mechanism	PMSA	Primary Mirror Segment Assembly
MIRI	Mid-Infrared Instrument	PMBSS	Primary Mirror Backplane Support Structure (BSF + BP)
MLI	Multi-Layer Insulation	Q	Heat
NASA	National Aeronautics and Space Administration	SI	Science Instrument
NGAS	Northrop Grumman Aerospace Systems	SINDA	Systems Improved Numerical Differential Analyzer modeling tool
NIRCam	Near-Infrared Camera Instrument	SM	Secondary Mirror
NIRSpec	Near-Infrared Spectrograph Instrument	SMA	Secondary Mirror Assembly
OA	Optical Assembly	SMSS	Secondary Mirror Support Structure
OGSE	Optical Ground Support Equipment, a series of pre-OTIS Optical pathfinder tests	SVTS	Space Vehicle thermal Simulator
OTE	Optical Telescope Element	TM	Tertiary Mirror
OTIS	Optical Telescope Element plus Integrated Science Instrument Module (OTE + ISIM)	TPF	Thermal Pathfinder test
PG	PhotoGrammetry cameras	W	Watt(s)



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